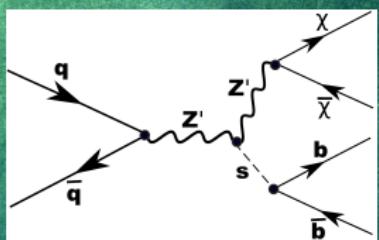
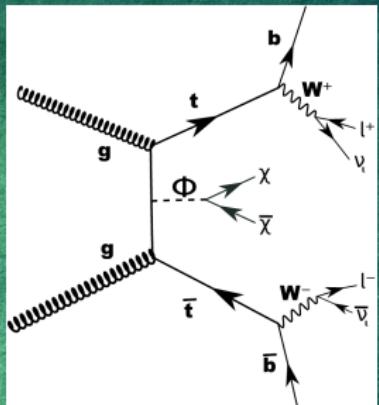


On the trail of dark matter by use of simplified models at CMS

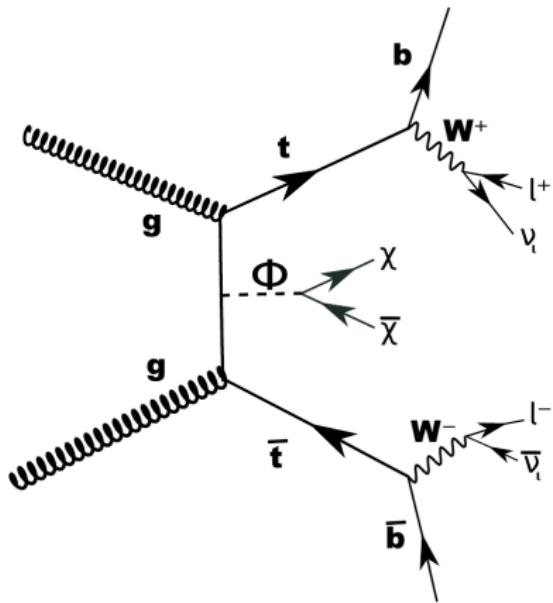
Michael Dürr, Alexander Grohsjean,
Kai Schmidt-Hoberg,
Christian Schwanenberger,
Nicole Stefanov

DPG-Frühjahrstagung Würzburg,
20th March 2018

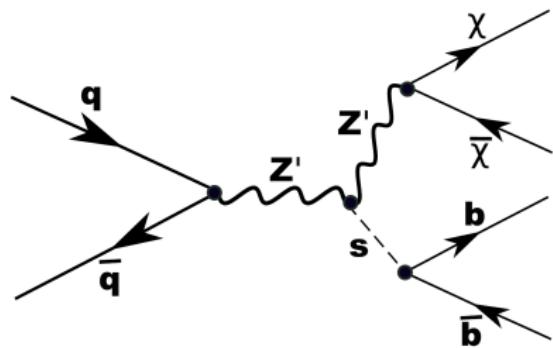


Outline

- Weakly Interacting Massive Particle hypothesis
- Search at LHC possible

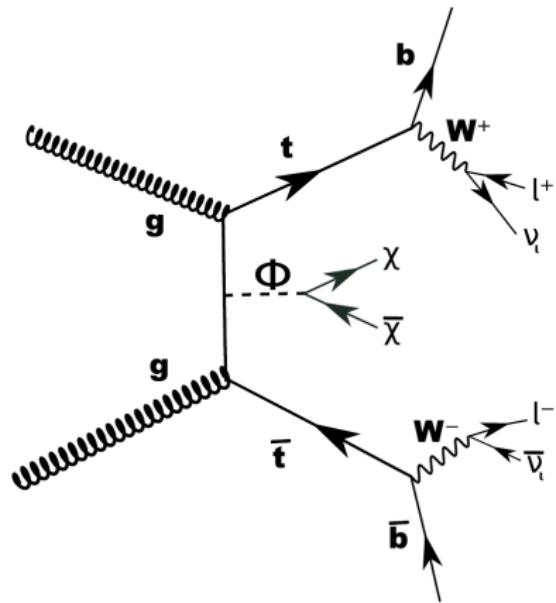


with Spin-0 mediator



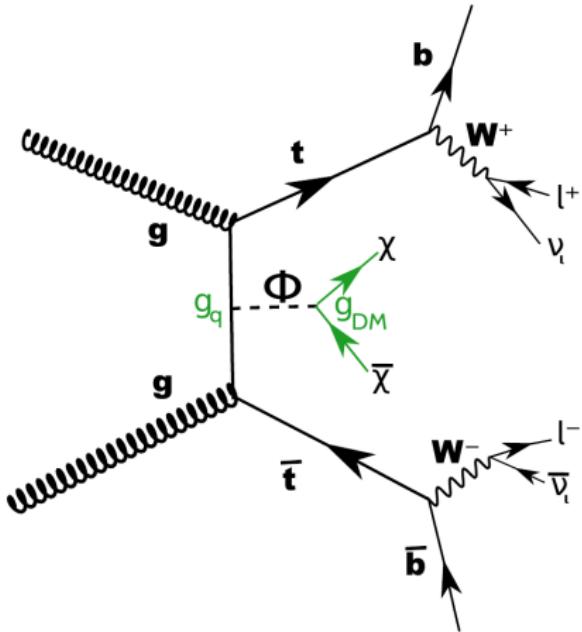
with Spin-0 and Spin-1 mediator

- Weakly Interacting Massive Particle hypothesis
- Search at LHC possible



with Spin-0 mediator

Search by use of simplified DM models



Framework

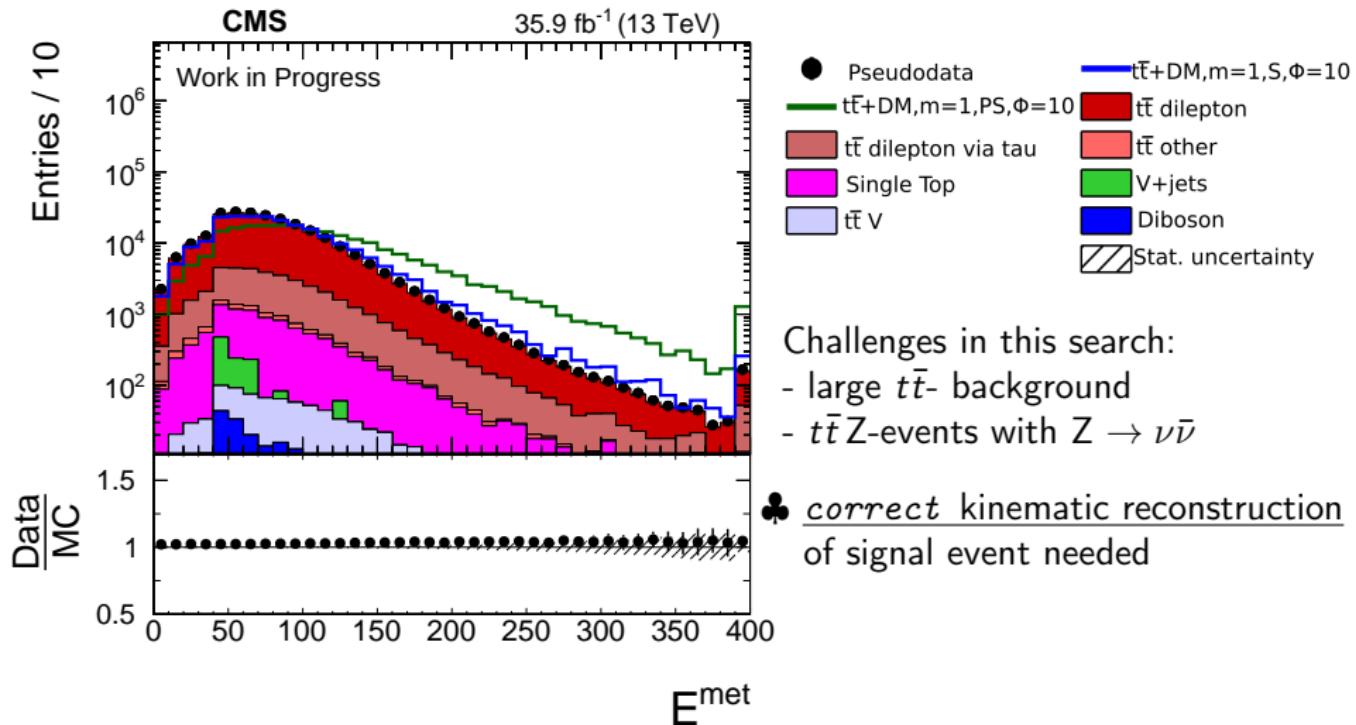
- DM Dirac fermions
- (pseudo-) scalar Φ
- Four parameters for description needed:
 $g_q, g_{DM}, m_\chi, m_\Phi$

Direct recoil of DM against SM particles

Motivation

- Yukawa-like coupling
⇒ strongest coupling to heaviest quark
(= top quark)
- test nature of mediator

Search by use of simplified DM models



Kinematic reconstruction

Sonnenschein's approach

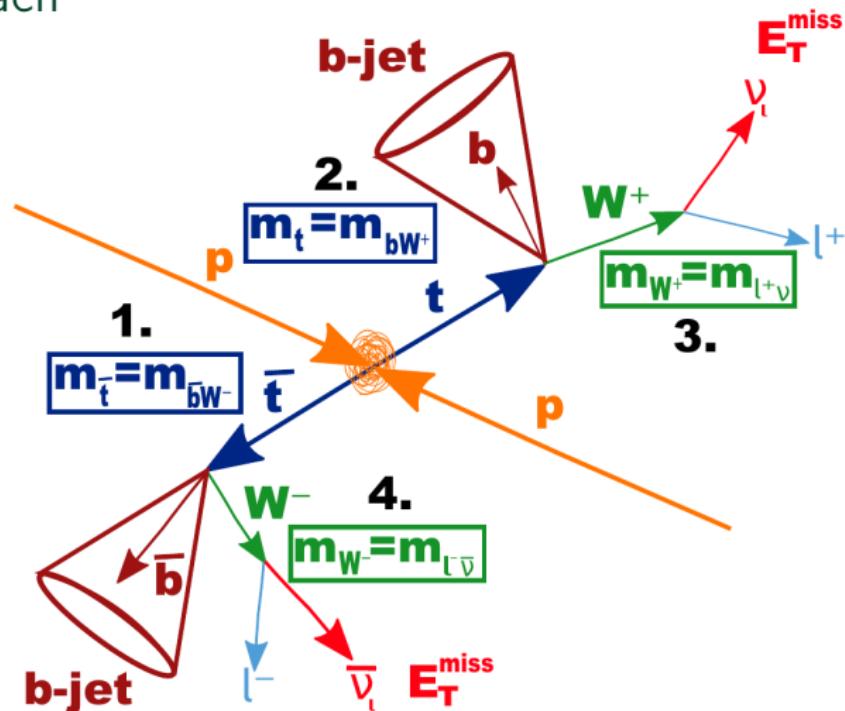
Unknown:
neutrinos' 4-momentum

⇒ 6 unknown variables
⇒ 6 constraints needed

$$5. E_{T,x}^{miss} = p_{\nu_x} + p_{\bar{\nu}_x} \quad \&$$

$$6. E_{T,y}^{miss} = p_{\nu_y} + p_{\bar{\nu}_y}$$

arXiv:0603011v3



⇒ results in a polynomial of $p_{\nu_x} \rightarrow$ in total max. 4 solutions

Kinematic reconstruction

Betchart's approach, arXiv:1305.1878v2

⇒ Splitting decay chain of top quark event into two steps
 $t \rightarrow W b$ and $W \rightarrow l \nu$

$$E_t = \underbrace{E_b + E_W}_{step1} = E_b + \underbrace{E_l + E_\nu}_{step2} \quad \& \quad \vec{p}_t = \vec{p}_b + \vec{p}_W = \vec{p}_b + \vec{p}_l + \vec{p}_\nu$$

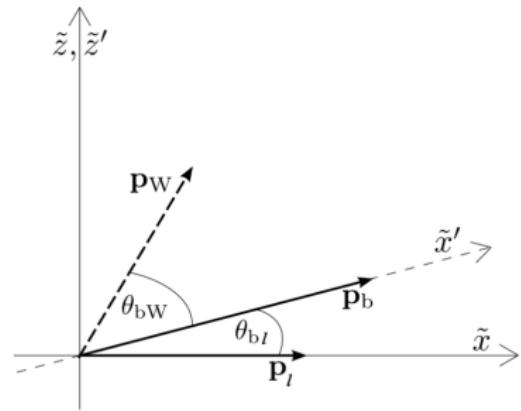
Kinematic reconstruction

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$$\begin{aligned} m_t^2 &= E_t^2 - \vec{p}_t^2 = (E_b + E_W)^2 - (\vec{p}_b + \vec{p}_W)^2 \\ &= m_b^2 + m_W^2 + 2E_b \cdot E_W - 2\vec{p}_b \cdot \vec{p}_W \cos \theta_{bW} \\ \bullet \vec{p}_W \text{ restricted to ellipsoid's surface} \end{aligned}$$



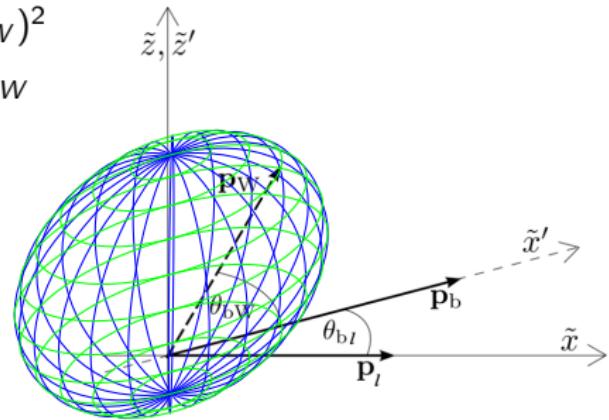
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Kinematic reconstruction

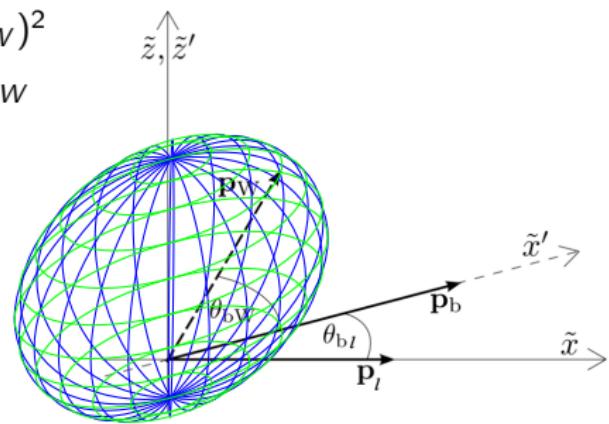
Betchart's approach, arXiv:1305.1878v2

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- \vec{p}_W restricted to ellipsoid's surface
- Solve step 2 for \vec{p}_ν analogously
+ shift by \vec{p}_l to get 2. ellipsoid for \vec{p}_W
- intersect ellipsoids



Kinematic reconstruction

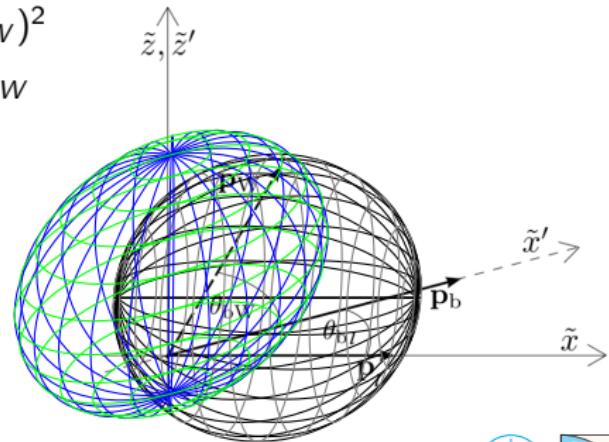
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Kinematic reconstruction

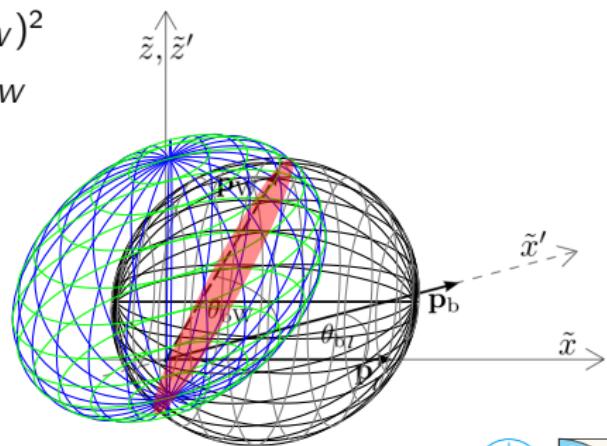
Betchart's approach, arXiv:1305.1878v2

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- intersect ellipsoids
⇒ ellipse with ∞ solution points \equiv ES



Kinematic reconstruction

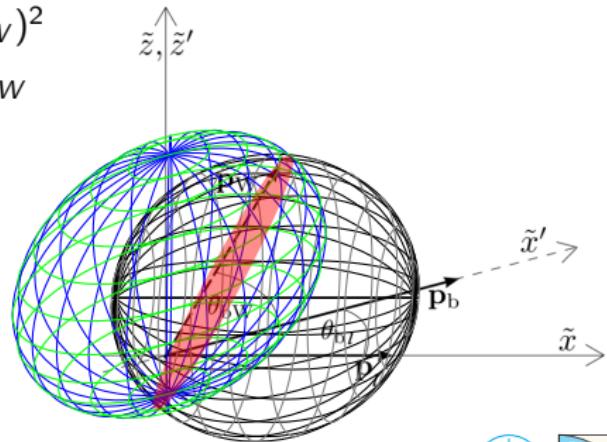
Betchart's approach, arXiv:1305.1878v2

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- \vec{p}_W restricted to ellipsoid's surface
- Solve step 2 for \vec{p}_ν analogously
+ shift by \vec{p}_l to get 2. ellipsoid for \vec{p}_W
- intersect ellipsoids
⇒ ellipse with ∞ solution points \equiv ES
- analogously for antitop quark
- intersect ES of \vec{p}_ν with MET ellipse
- in total max. 8 solutions



Kinematic reconstruction

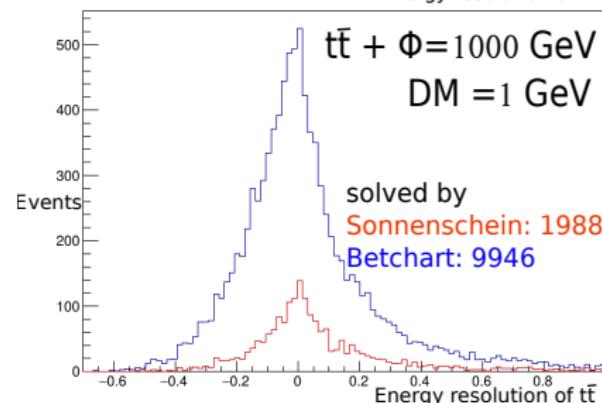
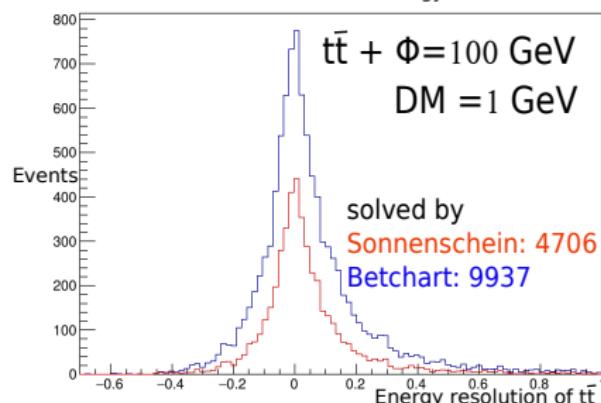
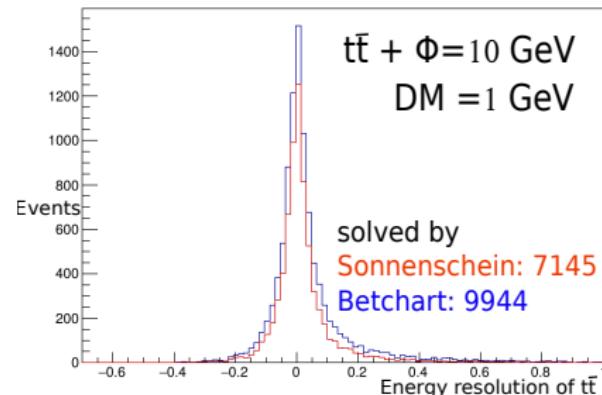
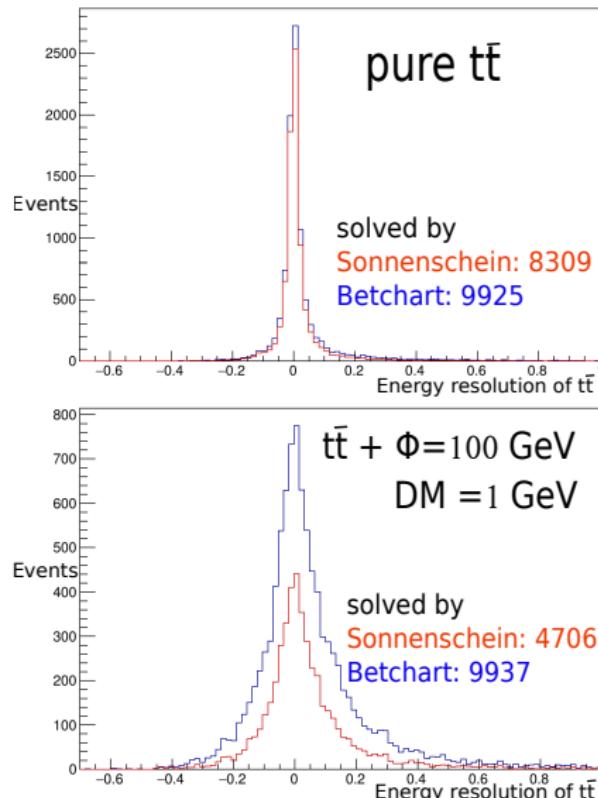
Signal events used for
Betchart vs. Sonnenschein comparison

Test on

- Monte Carlo samples on parton level
- With 10000 events, respectively
- Betchart and Sonnenschein algorithm provided
 - with fixed values for $m_t, m_{\bar{t}}, m_{W^-/+}$
 - and correct b-jet matching
- NLO model files of
<http://feynrules.irmp.ucl.ac.be/wiki/DMsimp>
used for generation of $t\bar{t}DM$ samples

Kinematic reconstruction

Betchart vs. Sonnenschein

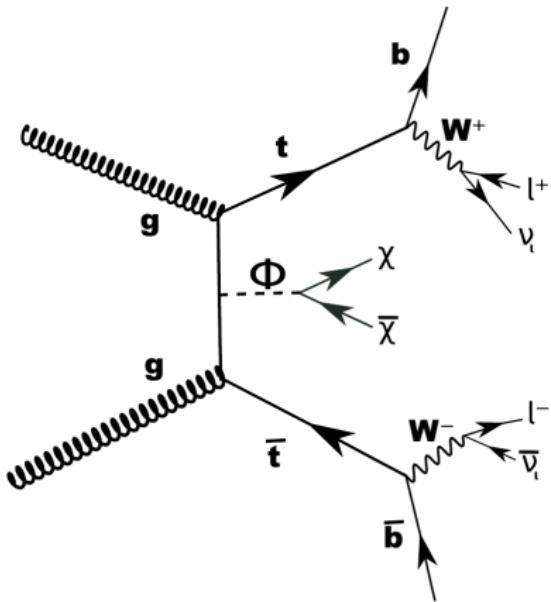


♣ With higher Φ mass, Sonnenschein fails increasingly, Betchart still stable

Nicole Stefanov - DPG-Frühjahrstagung, Würzburg, 20.03.2018 - 8/16



What's next?

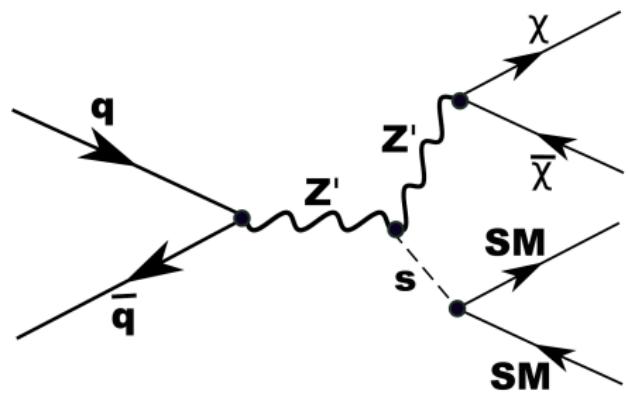


with 1 Spin-0 mediator

♣ Continuing with Betchart's approach

- Ellipse of solution = ES gotten without MET
→ more robust for higher Φ masses
- Finding condition without using MET for determining the correct solution on ellipse
→ Search gets more sensitive for higher mediator masses (> 100 GeV) where Sonnenschein breaks down

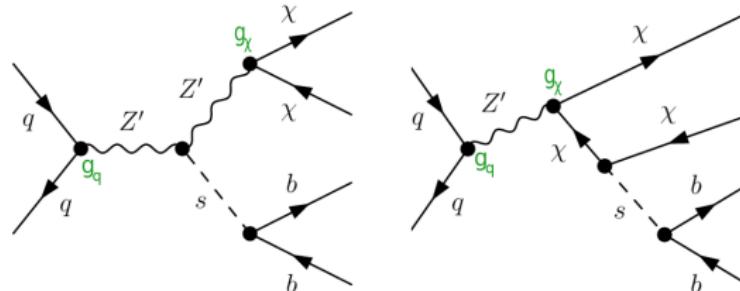
Extended simplified DM models



with 1 Spin-0 and 1 Spin-1 mediator

Extended simplified DM models

with Spin-0 mediator s and Spin-1 mediator Z'



Motivation DM mass generation needed +
relaxing experimental constraints from DM relic abundance

- new $U(1)'$ gauge group for dark sector → symmetry breaking
- gives rise to Z' and DM masses

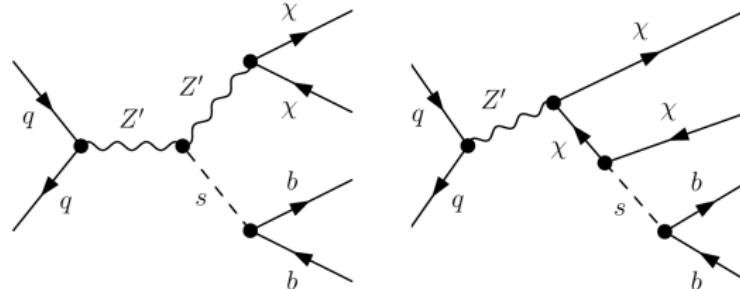
DM Majorana fermions

Six parameters for description needed: $g_q, g_\chi, m_\chi, m_{Z'}, m_s, \theta$
with θ = mixing angle between dark and SM Higgs

arXiv: 1510.02110, 1606.07609, 1701.08780

Extended simplified DM models

with Spin-0 mediator s and Spin-1 mediator Z'



↑ Recoil of DM particles ↑
against visibly decaying dark
Higgs

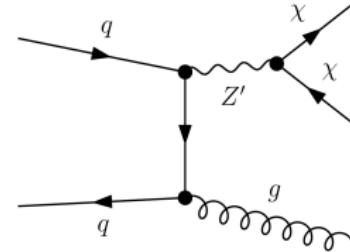
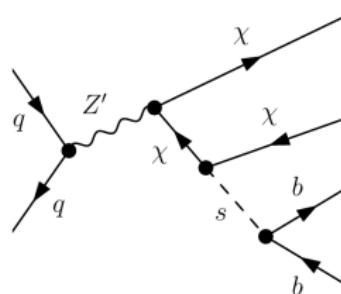
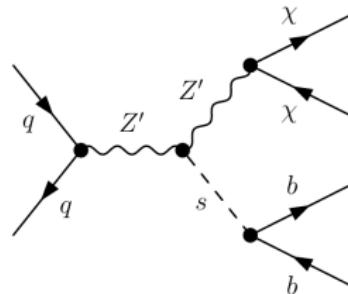
- ♣ For $m_s < m_\chi$, annihilation channel $\chi\chi \rightarrow ss$ facilitates reaching the correct DM relic abundance
⇒ Experimental constraints relaxed
- ♣ DM production via additional Z' mediator gives rise to dark-Higgs strahlung

arXiv: 1510.02110, 1606.07609, 1701.08780



Extended simplified DM models

with Spin-0 mediator s and Spin-1 mediator Z'



↑ Recoil of DM particles
against visibly decaying dark
Higgs

↑ Recoil of DM against SM
particles

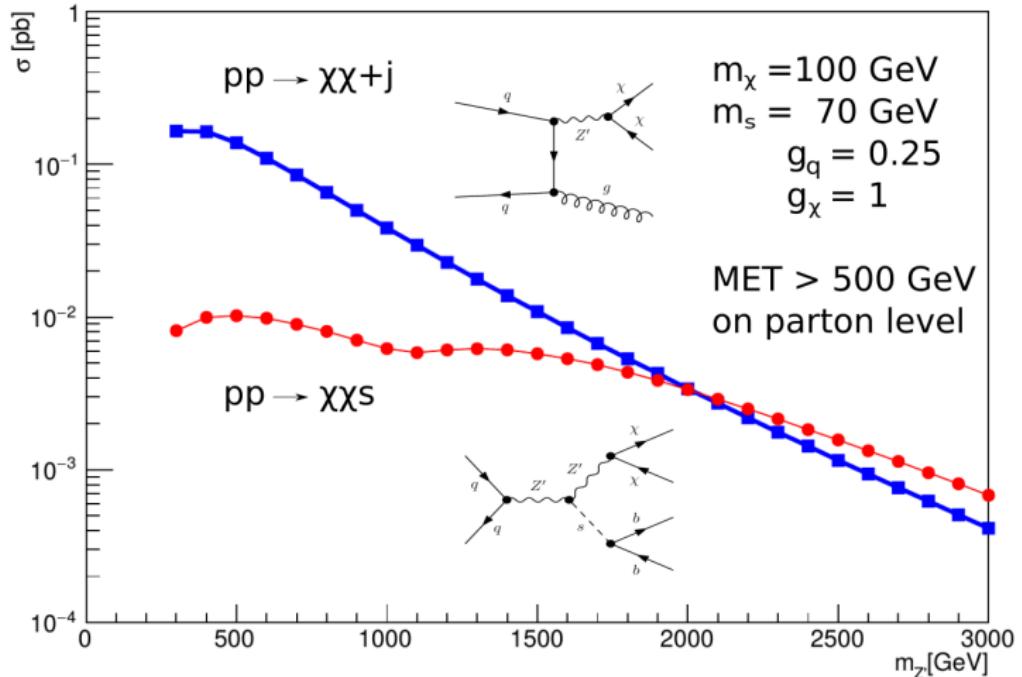
How does it compare to the mono-jet search?

arXiv: 1510.02110, 1606.07609, 1701.08780



Dark Higgs: New search possibilities

$pp \rightarrow \chi\chi + j$ vs. $pp \rightarrow \chi\chi s$



⇒ signature of
decaying
dark higgs
channel gains
in importance

- With higher Z' mass, especially \uparrow at $m_{Z'} \approx 1500$ GeV when
 $2 \cdot \sigma(pp \rightarrow \chi\chi s) \approx \sigma(\chi\chi + j)$

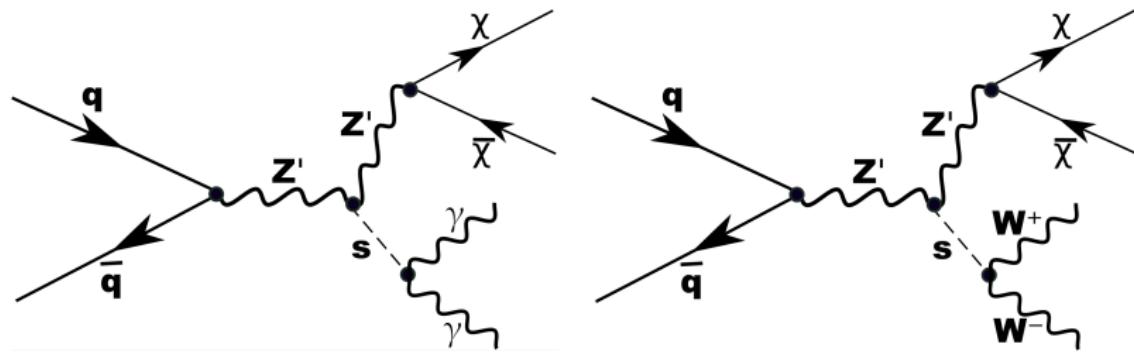
What's next?

$s \rightarrow b\bar{b}$ covered in arXiv: 1701.08780

⇒ see talk T 77.8 on Thursday, 22nd March, at 18.30
„Hunting the Dark Higgs at CMS“ by Samuel Baxter

But other decay modes also possible:

⇒ Study expected LHC sensitivity for various Z' masses for different decay modes:



Conclusion.

- **Simplified model with Spin-0 mediator**

Kinematic reconstruction by Betchart will increase sensitivity in search for $t\bar{t}\text{DM}$ considerably

- **Simplified model with Spin-0 and Spin-1 mediators**

Search for $p\ p \rightarrow \chi\chi s$ production will be extended by including further s decay modes

Thanks for your
attention!

Any
Questions?