Luminosity measurement at CMS

Jessica Leonard, DESY-Zeuthen On behalf of the CMS Collaboration



Luminosity at CMS overview



Luminosity calibration

Offline luminosity measurement: pixel cluster counting



Several effects need to be accounted for during the analysis.

The relationship between number of pixel clusters and luminosity can become nonlinear when pixel clusters belong to more than one track. However, the loss of linearity was determined to be very small, only 1% at a pileup value of 200 (below).

The detector can also be affected by "dynamic inefficiencies" when the data rate becomes very high, causing the data acquisition system to become busy and preventing data taking. This effect is in general quite small, less than 0.5% overall (upper right).

"Afterglow" due to detector material activation can cause out-of-time response in the pixel cluster count. The afterglow effect was modeled assuming an exponential decay, and the effect was determined to be $\sim 2\%$.

In general, the uncertainty of the luminosity integration is low, contributing a total of 1.2% on the overall uncertainty, which is 2.5%.



0.5

0.5

2

0.3

0.2

0.5

0.2

0.1

0.5

0.5

2.5

Stability of measurement

The luminosity calibration constant, σ_{vis} , should remain constant in time, but this is only true of an ideal luminometer. Since it depends on the acceptance and efficiency of the luminometer, it may need to be corrected according to the measured acceptance and efficiencies. Using only channels that have been active for the full run mitigates this need to a certain extent, but there are many other effects that can play a role.

One method of ensuring σ_{vire} remains constant is to check the pixel layers against each other. This gives excellent results, as shown in the figure below. The relative comparisons show stability at the 0.5% level.





CMS silicon pixel detector

- 3 barrel layers, 2 endcap disks per side - 100 x 150 µm sensors - 66 million channels, 96.3% always alive - <0.1% occupancy at 10³⁴ cm⁻²s⁻¹ - Max readout rate 100kHz

The pixel detector is especially suited to luminosity measurement because of its high fraction of always-alive channels, minimizing variation in detector acceptance, and its very low occupancy at high luminosity, which means count rates are linear with luminosity. It can only operate during stable running conditions, which means it must only be used as an offline luminometer. However, the system is stable and precise and, as such, is the reference luminometer for CMS.



Another stability check is a comparison with the Z->µµ cross section. By definition the true Z->µµ cross section is constant with time and constant beam energy. Therefore, if a different cross section is measured with time, it can be corrected and the corrections also applied to the luminosity measurement. This study is in progress.

Online luminosity measurement: forward hadronic calorimeter





The HF luminosity measurement uses the zero-counting technique. The number of bunch crossings with zero hits is counted. The negative logarithm of this quantity is proportional to the number of interactions in the same time period, due to the Poissonian behavior of the hit rate. When the fraction of zero hits is very high, the relationship between luminosity and the measured quantity is linear.

HF also participated in Van der Meer scans for absolute luminosity calibration. In the examples shown below, the rate vs beam separation curves for individual bunches were fitted with a double-





Upgrades to online luminometers

After the upgrade, the online luminosity measurement system will include several more subsystems. This will introduce redundancy into the system, ensuring continuous performance in case one subsystem drops out. In addition, having multiple measurements provides confirmation of the measured luminosity value.

The Fast Beam Condition Monitor (BCM1F) will consist of 24 single-crystal diamonds situated in two parallel planes on either side of the interaction point. The Pixel Luminosity Telescope (PLT) will consist of 16 3-layer silicon telescopes, situated just outside BCM1F in the z-direction. The subsystems will each

